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TECHNICAL NOTE Monitoring the Magmas Fuelling Volcanic Eruptions in Near-real-time Using X-ray Micro-computed Tomography

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A novel application of X-ray micro-computed tomography is described, which can be used to rapidly characterize chemical populations of natural olivine crystals in erupted basalts. This technique can be deployed during volcanic crises to directly track changes in magma components of an erupting system in near-real-time. Such changes are fundamental in controlling eruption style, duration and intensity. We demonstrate a method that can generate data from hundreds of crystals within hours, which allows time-series petrological data to be recorded and interpreted alongside various complementary monitoring techniques (e.g. seismicity, ground deformation). Our direct-detection will allow greater understanding of the dynamics of sub-volcanic magma plumbing systems, and can provide important insights into how an eruption may evolve. The same technique can also be used to generate rich baseline datasets from eruption sequences in the geological record in a more efficient manner than conventional methods allow.

KEY WORDS: X-ray micro-computed tomography; olivine composition; volcanic eruption monitoring

INTRODUCTION

Volcanic eruptions are often triggered, sustained, or terminated by changes in magmatic conditions that affect the physical behaviour of the magma; these include temperature (T), pressure (P), melt composition including volatile (e.g. H_2O) content (X), and consequent gas and mineral phase stability (Bardintzeff & Bonin, 1987; Cashman, 1992; Papale et al., 1998; Pinkerton et al., 2002; Rutherford & Devine, 2003; Izbekov et al., 2004). Mixing between different magma batches is frequently implicated as the cause of such changes, with a common scenario being a new magma pulse entering the sub-volcanic plumbing system (Sparks et al., 1977; Huppert et al., 1982; Sigmundsson et al., 2010). Developing the capability to detect such changes, understand their influence, and ultimately forecast future eruptive events with high temporal resolution is a continuing goal of volcanologists (Aiuppa et al., 2002; Sparks et al., 2012), owing to the enormous capacity for volcanic

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eruptions to cause both local (Tanguy, 1994) and global natural disasters (Self, 2006).

In addition to predicting the beginning of an eruption, other important questions are: 'How large, for how long a duration, and therefore how hazardous could a new eruption become?' These forecasts are critical to effective planning and crisis management. For instance, in an Icelandic setting, an initial small fissure eruption may have the potential to evolve into a small magnitude 2010 Eyjafjallajökull-style event (Stevenson *et al.*, 2012), or a large magnitude eruption such as the AD 1783–1784 Laki event (Thordarson & Self, 2003). The frequency of explosive eruptions within Iceland and in other volcanic settings around the world in historical times, and the large variability in scale they exhibit (Thordarson & Larsen, 2007), emphasizes the importance of these questions.

The role of igneous petrology in understanding volcanic risk

Igneous petrology provides insights into the conditions of magma generation and storage and the drivers of volcanic processes. The saturation of phenocryst minerals during the evolution of a magma is controlled by variations in P, T and X of the host melt through time (Wilson, 1989). Cooling magmatic systems typically evolve from mafic (Mg-rich) to more felsic (SiO₂-rich) compositions in response to fractionation of mafic minerals, and the Mg# $[100 \times Mg/(Mg + Fe)]$ of the residual liquid decreases with time (Villiger et al., 2004). Important physicochemical changes in the magma occur simultaneously with this differentiation, including increases in volatile concentrations and changes in crystal type, abundance and chemistry. Concomitant changes in magma viscosity also occur (Giordano et al., 2008). The erupted volcanic products may thus preserve a record of the changing magmatic conditions in P-T-X space with time in the chemistry of the phenocryst minerals (Fram & Lesher, 1997; Thy et al., 2006).

Interaction between batches of magma with different intensive parameters (including P, T, X, crystal and gas bubble content) is a common process (Streck, 2008) that can add further complexity to the overall magmatic system as it evolves. Because eruption dynamics are largely controlled by the physicochemical state of the magma (Papale *et al.*, 1998), monitoring such interaction is of considerable importance. Directly observable consequences of significant magmatic change induced by mixing processes include changes in crystal populations and in crystal chemistry (Sigmarsson *et al.*, 2011).

Magmatic provenance and change indicated by olivine crystal chemical populations

The composition of olivine is particularly sensitive to changes in magma composition, P, T and oxygen fugacity (Putirka, 2008; Matzen *et al.*, 2011). In basalt-dominated

terrains such as Iceland, the olivine chemical population(s) in the erupted basalts can be used to identify the likely magma source regions (e.g. deep versus shallow crust) and thus differentiate between a large, deep-seated rift event or a small, shallow, remobilized magma event (Sigmarsson *et al.*, 2011). As such, olivine chemical populations could provide an indication of the eventual scale and progression of a new eruption. Furthermore, changing crystal chemical populations throughout an eruption sequence may indicate changing magmatic conditions, potentially owing to changes in the relative volumes of interacting magma batches each contributing the signature of their crystal load to the volcanic products.

Olivine is primarily a solid-solution between forsterite $(Mg_2SiO_4: denoted as Fo_{100} \text{ or } Mg\# = 100)$ and fayalite $(Fe_2SiO_4: Fo_0, Mg\# = 0)$, and the Fe–Mg mineral–melt partitioning is well understood (Roeder & Emslie, 1970). Therefore the Mg# of olivine precipitating from a melt represents a proxy for the stage of differentiation of the equilibrium melt. Hence olivine Mg# populations can provide important constraints upon local melt conditions and larger-scale magmatic processes (Thomson & Maclennan, 2013).

Whereas the rim of a growing magmatic olivine crystal reflects the composition of the adjacent melt, its interior provides a record of earlier conditions specific to that crystal. A number of processes contribute to the crystal population in a magma. Phenocrysts grow in the magma containing them, antecrysts are earlier precipitates from a discrete but kindred magma/cumulate mush, and xenocrysts are plucked from unrelated wall-rocks or cumulate mush (Hildreth & Wilson, 2007; Streck, 2008). Natural magmatic systems can therefore be highly complex, and involve interaction between multiple magma batches, each bringing their distinct crystal population and history to the mix.

This rich complexity can provide powerful insights into magmatic systems if the contributions from individual components can be identified (Kahl *et al.*, 2011). If sufficient numbers of crystals can be analysed to produce statistically robust datasets (Thomson & Maclennan, 2013), changes to magmas and/or contributions from different magma batches can be tracked in the erupted crystals throughout an eruption sequence. In this manner, monitoring relative changes to crystal chemical populations in near-real-time during a volcanic eruption would represent a major advance.

THE PROBLEM

Conventional chemical analysis of crystals requires time intensive preparation. Each mount may take several hours to prepare (to fully set the resin, which is then ground to expose the grains, and polished to a high-quality finish before being carbon coated). Crystals can be analysed only one at a time by electron-probe micro-analysis (EPMA). A few minutes per analytical spot are sufficient to produce quantitative data for major elements and a small number of trace elements. Optimized EPMA and fully quantified energy-dispersive X-ray spectroscopy system set-ups may increase sample throughput, yet necessarily sacrifice precision and the number of elements analysed. Even with an optimized system involving electron beam analysis the sample material still needs to be mounted in resin, polished and carbon coated. Although direct calls have recently been made to monitor eruptive material (Sigmundsson & Hoskuldsson, 2010), and despite the long-recognized advantages of applying petrological data to volcano monitoring (Devine et al., 1998a, 1998b), rapid high-volume throughput in a time-sensitive situation (i.e. a current eruption) using standard methods remains problematic.

OUR APPROACH

To overcome the above impediment we turn our attention to a far faster method, X-ray micro-computed tomography (XMT; Stock, 1999), and use it to analyse large numbers of entire crystals simultaneously. XMT is a non-destructive density-sensitive imaging method that allows density contrasts to be mapped in three dimensions. Samples are rotated in an X-ray beam while a high-resolution scintillator fronted CCD detector collects images (projections) at different angles. The projections are then 'reconstructed' using numerical algorithms to produce a full 3D render of the internal structure of the object (Stock, 1999). Modern laboratory and synchrotron imaging systems can acquire data in as little as a few seconds, at a voxel (the 3D equivalent of a pixel) resolution of $<2 \,\mu m$. Applications to geological materials prior to our study have focused upon identifying different phases and minerals, and investigating their relationships in static or dynamic three dimensions (Cnudde & Boone, 2013). Greyscale values can be compared between scans if secondary standards are used that encompass the X-ray attenuation coefficient of the material of interest; these allow correction for variations in absolute greyscale value that can be caused by fluctuations in X-ray beam current and flux (owing to filament age, instrument behaviour, etc.).

The X-ray linear attenuation coefficient of forsterite and fayalite differs significantly under typical laboratory scan conditions (40-60 kV, Fig. l). Here we investigate (l) whether this difference in attenuation coefficient is sufficient for useful discrimination between olivine crystal compositions, (2) whether observed differences can be confidently mapped to measured composition differences, and (3) how quickly these data can be generated. We have simulated the throughput of samples in a time-sensitive situation to test how quickly volcano observatories might be able to produce and interpret data during a volcanic



Fig. 1. The dependence of linear attenuation coefficient on X-ray energy for olivine. The linear attenuation coefficient increases with decreasing Mg#. Scanning at lower energies, where the contrast in attenuation coefficient is greatest, results in a wider dynamic range in the reconstructed 3D images. The attenuation coefficient curves for olivine compositions bracketing our samples are shown. The Nikon XTH 225ST XMT system was run with a polychromatic X-ray beam with a maximum energy of 65 kV, and a peak flux at ~40 kV.

crisis. We used a sample of tephra erupted from the 2010 Fimmvörðuháls flank eruption prior to the Eyjafjallajökull summit eruption in the Eastern Volcanic Zone (EVZ) of Iceland, as similar material is most likely to be sequentially sampled during future eruptions.

METHOD

The tephra sample was manually crushed using a steel rolling pin and an aluminum bash-plate before being sieved. One hundred grains containing an olivine crystal were picked directly from the 0.5-2.0 mm fraction (grains containing whole, intact crystals were preferred), loaded into ~ 2.0 cm lengths of plastic straw (internal diameter = 4 mm), and capped with pre-prepared resin discs. Three grains containing plagioclase crystals of a similar size to the olivine were added as a control. This experimental charge was then scanned using a Nikon XTH 225ST XMT system at the University of Manchester at a maximum energy of 65 kV using a 10 Al filter to minimize beam hardening. This system has a standard static X-ray source and detector (which resolves 2000×2000 pixels); thus with decreasing distance between the sample and the source, the greater the absolute resolution. The sample rotates around a vertical axis while projections are taken about 360°. Using more projections decreases the number and intensity of reconstruction artefacts, at the expense of time. We found that the balance between data quality and speed of acquisition was optimal at 2000 projections (every 0.18°) per scan.

The data were reconstructed using proprietary software to produce a 3D volume containing all 103 grains with a voxel resolution of $\sim 2.5 \,\mu$ m. Digital image processing techniques were used to isolate each olivine crystal using *Aviso®* and *Fiji*. This involved using pre-existing and freely available plugins to 'train' the software to distinguish olivine from plagioclase, glass and air. Once identified, each crystal in the olivine fraction was digitally eroded inwards from the rim by the equivalent of ~15 μ m to ensure only the core of the crystals remained. This avoided the sampling of heterogeneous rims, with the attendant potential for error, during subsequent analysis. The average greyscale value of a 'spot' of identical size (25 μ m radius) was then calculated for each crystal core, excluding obvious inclusions, cracks or zoning. This last step was performed manually to ensure a robust dataset with which to compare against future, more sophisticated/automated, image processing methods, and to validate against conventional methods whereby analysis positions are chosen manually.

Standard EPMA techniques were used to measure the compositions of 100 olivine crystal cores from a second split of the same sample using a JEOL 8230 electron microprobe at the University of Leeds (see Appendix for full analytical conditions).

RESULTS

The entire XMT method took <2.5 h (see Fig. 2). Total sample preparation time was ~ 20 min. The scanning and reconstruction steps (image acquisition) were performed in ~ 30 min each. Image processing (including manual selection) took ~ 60 min.

In the studied sample olivine, plagioclase, glass and air that make up the 3D volume were easily distinguished based on habit, texture and greyscale value. The plagioclase (used as a control) was identified by its habit, texture and (low) greyscale value. Thirteen olivine crystals were discarded owing to cracks or inclusions. The reconstructed volume contains olivine crystals with a range of greyscale values, and often reveals zoning around homogeneous cores (Fig. 2). Thus we can demonstrate that XMT is capable of resolving even subtle changes in Mg# in natural olivine crystals.

A frequency diagram of the greyscale values was constructed from both the XMT-spot (n = 87) and EPMAspot (n = 100) datasets using nine bins $(\sim 10\%$ of each n)of equal range (Fig. 3). A distinct pattern of greyscale value population frequency is observed, and is matched by the compositional variations determined by EPMA (Fig. 3). Subtle differences between the two datasets are due to a combination of the non-linear energy dependence of the X-ray attenuation coefficient, which is more pronounced at high Mg# (see Fig. 1), and the use of two sets of olivine crystals from the same sample. The EPMA data are reported in Table Al and the XMT data in Table A2. Rank-order and linear-array plots of the XMT and EPMA data are provided along with the raw data in the Appendix (Fig. Al and Table A3, respectively). The range of greyscale values in each XMT data bin is provided in Table A4.

Identification and quantification of chemical population(s) of macrocryst phases is an important step in unravelling the petrogenesis of igneous rocks, and is a key result of our study. Figure 3 demonstrates the existence of a dominant olivine chemical population of \sim Fo₈₀ in the studied sample. Two sub-dominant populations are also evident, at \sim Fo₇₆ and \sim Fo₈₄, which indicates disequilibria of the system at the time of eruption. Taken at face value, these three peaks may demonstrate the presence of a number of magmatic components in the erupted sample. These have arguably mixed shortly before eruption, as olivine not in equilibrium with the host melt will be eventually resorbed.

We find that the XMT method is far faster than standard analysis techniques. It has the ability to accurately distinguish chemical populations, as confirmed by the EPMA data. The use of secondary standards that bracket and intersperse the X-ray attenuation coefficient range of olivine should allow fully comparable data between scans. We conclude that XMT holds great potential to be used as a high-throughput, rapid and accurate method for characterizing crystal chemical populations, which opens up a powerful new approach to volcanic eruption monitoring, and will allow the generation of rich baseline datasets.

FUTURE APPLICATIONS

To make best use of olivine XMT data in a time-critical scenario such as a volcanic eruption, the absolute chemical composition and zonation of olivine crystals must be assessed soon after eruption. Optimized EPMA techniques on polished crystal mounts produce high-precision data, but they are time intensive, destructive and limited to 1D and 2D profiles. With appropriate calibration and use of secondary standards, our XMT technique could approach the precision and accuracy of current EPMA methods in terms of major element chemical composition, and open further avenues for its application in volcanic eruption monitoring.

Recently developed understanding of solid-state Fe-Mg diffusion between zones of different Mg# within olivine has effectively turned zonation profiles in each crystal into a stopwatch for magmatic processes (Costa & Dungan, 2005; Morgan & Blake, 2006; Dohmen & Chakraborty, 2007; Costa et al., 2008). Characterization of the zoning of single olivine crystals and of crystal populations in eruption sequences can enhance our understanding of the evolution of sub-volcanic magma plumbing systems through time, on a range of scales (Kahl et al., 2011, 2013). With large enough olivine composition datasets, a detailed picture of the magmatic processes fuelling basaltic eruptions through time can be obtained, and compared with seismic unrest and ground deformation signals. This information can aid interpretation of geophysical unrest signals during pre- and syn-eruption events in the future.



Fig. 2. XMT workflow for determining olivine crystal composition populations within 2:5 h of sample receipt. Crystals are easily distinguished on the basis of texture and shape. Low-density olivine (high Mg#) appears darker than higher density olivine (low Mg#). Refinement of the experimental setup may allow faster scanning, and trainable algorithm-based analysis may decrease this workflow time further without sacrificing data quality. Volumes for greyscale value averaging were selected to simulate an EPMA spot, and were placed within the crystal cores, shown schematically as white spots.



Fig. 3. Comparison between XMT greyscale value data and EPMA (Mg#) data from olivine crystals in tephra erupted during the 2010 Fimmvörðuháls eruption. The XMT data were generated using the XMT workflow described in the text and summarized in Fig. 2. A single EPMA spot was analysed in each crystal core (see the Appendix for analytical details). The values in each dataset were divided into nine bins of equal range. The peaks in each population correlate well.

Within basaltic systems such as the active EVZ of Iceland olivine is almost ubiquitous, and records a variety of origins (Passmore et al., 2012). Olivine in these rocks is commonly chemically zoned: generally a homogeneous core is fringed by comparatively narrow zones of different olivine composition (Thomson & Maclennan, 2013). Volcanic systems such as these represent prime candidates for XMT-based generation of rich datasets of olivine composition populations and magmatic timescales through eruption sequences. Furthermore, integrating crystal size distribution (CSD) and modal abundance data with our rapid composition determination approach could be a powerful tool in understanding magma dynamics and, by extension, understanding crystal-melt relationships. XMT can be used to measure CSDs (Jerram et al., 2009), and each crystal can be assigned a unique digital label, allowing large populations of crystals to be characterized by size and composition.

In the short term, linking olivine chemical population changes with well-resolved eruption phenomenology (duration, intensity, style) in a forensic manner and merging



Crystal core Mg#

Fig. 4. (a) Schematic illustration of hypothetical volcano behaviour (observed in real-time) to illustrate the response to major changes within the magma plumbing system during a protracted eruption. Light shades indicate silicic melts and Fe-rich olivine compositions (i.e. shallow equilibrated, evolved magma), whereas dark shades indicate Mg-rich compositions (i.e. deeply equilibrated, primitive magma). (b) Olivine crystal chemical population frequency curves stacked through time, corresponding to the eruption sequences shown in (a). Such data could be generated from erupted material using a workflow similar to that described in this study, allowing insight into magmatic processes at a high temporal resolution. These data may be compared with the record of event chronology during volcanic eruptions to identify possible correlations with seismic data, ground deformation changes and gas monitoring data in near-realtime. Similar plots based on data from past eruptions may be used as empirical guides to future eruptive events.

this with geophysical datasets (Kahl *et al.*, 2013) holds potential to discover important patterns that could be recognized during a future eruption. There is an urgent need for well-documented analysis of past eruptive events, as they hold the key to well-constrained pattern recognition. These patterns could be used as empirical lead-indicators of eruptive change, duration and cessation, which are key issues during a volcanic crisis (see Fig. 4). Building this knowledge could be accelerated by using the highthroughput XMT method described here.

Although many hazardous volcanoes erupt olivinephyric basalts, many do not. The XMT method is not limited to olivine, for the principles of variable X-ray attenuation coefficient as a proxy for variable composition also applies to many other phenocryst minerals and glasses. Those with solid-solutions whose end-member compositions represent a large density contrast are particularly suited to this method, such as enstatite-ferrosilite and anorthite-albite. By applying a similar method to other minerals such as orthopyroxene and plagioclase, the XMT technique will be of use in a greater number of volcanic settings. In addition, XMT techniques are not limited to the study of volcanic samples; 3D applications for observing compositional variations in intrusive magmatic rocks and even mantle rocks (xenoliths; orogenic peridotites) could hold great potential. Since XMT generates high quality 3D textural data, it holds great potential for use as a provider of specific and timely particle shape, size and density information as inputs to next generation ash dispersion models.

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Fig. A1. Comparison between olivine chemical population data generated by EPMA and XMT from two splits of the same sample of Fimmvörðuháls tephra. (a) Decreasing rank-order plots of 100 EPMA Mg# spots (one in each crystal core) with increasing Mg#, and 87 XMT greyscale value spots (averaged 16-bit greyscale value) with increasing greyscale value. (b) Linear-array plots of the same EPMA and XMT data. The differential captures the degree to which each data point (each crystal) departs from a straight line between the highest and lowest value on the corresponding rank-order plot, so that comparisons are easier to make visually. It should be noted that higher greyscale values denote brighter voxels = lower Mg#. As such, the XMT linear-array order is inverted to allow comparison against Mg#.

Crystal	Oxide (wt %)									Mg#
	SiO ₂	TiO ₂	Al ₂ O ₃	FeO	MnO	MgO	CaO	NiO	Total	
FMVDCTA1a	40.04	0.02	0.02	12.18	0.15	46·71	0.24	0.38	99·77	87·10
FMVDCTB1	39.76	0.01	0.03	14·56	0.20	44.68	0.53	0.36	99.84	84·37
FMVDCT_C1	38.47	0.02	0.02	21.32	0.31	38.95	0.29	0.14	99.55	76·25
FMVDCT_D1	38.81	0.02	0.04	19.73	0.31	40.61	0.58	0.07	99.87	78·31
FMVDCT_E1	39.43	0.01	0.04	16.59	0.26	42·76	0.24	0.25	99.58	81·90
FMVDCT_F1	38.62	0.04	0.02	20.95	0.31	39.45	0.30	0.14	99.82	76·79
FMVDCT_G1	38·97	0.03	0.03	18·78	0.29	41.32	0.27	0.20	99.89	79·43
FMVDCT_H1	38.80	0.02	0.04	19.42	0.27	40.60	0.30	0.16	99·61	78·61
FMVDCT_11	38.56	0.02	0.04	20.54	0.33	39.71	0.58	0.10	99.58	77·23
FMVDCT_J1	39.67	0.02	0.02	15·26	0.24	44·18	0.27	0.26	99.93	83·56
FMVDCT_A2	39.40	0.01	0.04	16.08	0.22	43.36	0.24	0.26	99.62	82·59
FMVDCTB2	38·79	0.03	0.03	19.34	0.29	40.85	0.58	0.12	99.73	78·77
FMVDCT_C2	39.82	0.01	0.02	14·17	0.21	45·12	0.53	0.29	99.90	84·83
FMVDCT_D2	39.77	0.02	0.04	13·87	0.18	45·27	0.26	0.34	99.74	85·17
FMVDCT_E2	38·52	0.03	0.03	20.87	0.33	39.59	0.27	0.13	99.77	76·89
FMVDCT_F2	39.07	0.02	0.02	17.89	0.25	42·00	0.25	0.24	99.73	80.49
FMVDCTG2	39.09	0.01	0.02	18·15	0.27	41.72	0.24	0.20	99.72	80·16
FMVDCT_H2	39.11	0.02	0.04	18·51	0.28	41·52	0.25	0.19	99.92	79·75
FMVDCT_12	38.40	0.02	0.03	20.92	0.30	39.41	0.29	0.05	99·41	76·80
FMVDCT_J2	39.28	0.02	0.03	17.22	0.27	42.62	0.26	0.24	99.94	81·28
FMVDCTA3a	39.06	0.02	0.04	18·45	0.27	41.34	0.26	0.13	99·56	79·74
FMVDCTB3	39.40	0.01	0.06	16.77	0.26	42·85	0.26	0.23	99.84	81·78
FMVDCT_C3	38·31	0.03	0.03	22.33	0.35	38.23	0.58	0.07	99.63	75·02
FMVDCTD3	38·92	0.03	0.03	19.34	0.30	40.63	0.27	0.13	99.65	78·66
FMVDCT_E3	38·55	0.01	0.03	21.53	0.29	38.75	0.30	0.05	99.50	76·00
FMVDCT_F3	39.71	0.01	0.04	13.67	0.19	45·42	0.22	0.42	99.69	85·38
FMVDCTG3	38·37	0.02	0.02	21.52	0.38	39.05	0.27	0.16	99.78	76·07
FMVDCT_H3	39.20	0.03	0.06	17·91	0.27	41.83	0.18	0.17	99.66	80.40
FMVDCT13	38·97	0.03	0.04	18·56	0.30	41.45	0.29	0.18	99.82	79·67
FMVDCT_J3	39.26	0.04	0.06	18·20	0.28	41.79	0.27	0.16	100.06	80.13
FMVDCTA4a	39.60	0.02	0.03	14·98	0.21	44·35	0.24	0.28	99.70	83·88
FMVDCTB4	39·21	0.02	0.03	18·22	0.26	41.74	0.26	0.15	99.89	80.11
FMVDCT_C4	38.99	0.03	0.03	18·34	0.28	41.59	0.27	0.10	99.64	79·92
FMVDCTD4	39.35	0.03	0.04	17.00	0.26	42.67	0.25	0.17	99.77	81·50
FMVDCTE4	39.74	0.01	0.06	13.89	0.22	45·11	0.24	0.30	99.57	85·08
FMVDCT_F4	39.49	0.01	0.04	14·57	0.20	44·50	0.20	0.26	99.27	84·31
FMVDCTG4	39.19	0.02	0.03	17.77	0.27	41.91	0.27	0.18	99.65	80.55
FMVDCT_I4	39.22	0.02	0.03	17.25	0.28	42·73	0.26	0.22	100.02	81·29
FMVDCT_H4	39.12	0.01	0.03	18·12	0.27	41.75	0.27	0.14	99.71	80.19
FMVDCT_J4	38.86	0.02	0.03	18·92	0.29	41.06	0.58	0.14	99.59	79·21
FMVDCTA5	38.41	0.02	0.03	21.66	0.35	39.02	0.58	0.10	99.89	75·96
FMVDCTB5	40.02	0.01	0.02	12.35	0.18	46·23	0.22	0.42	99.48	86·81
FMVDCTC5a	40.11	0.00	0.06	12·75	0.18	46·15	0.23	0.35	99.84	86·41
FMVDCTD5	39·25	0.02	0.04	17.07	0.24	42.46	0.25	0.24	99.56	81·38
FMVDCT E5	39.23	0.02	0.03	17.31	0.27	42.44	0.28	0.17	99·76	81.14
FMVDCTF5a	39.26	0.02	0.03	17.04	0.30	42.68	0.25	0.17	99.76	81.43
FMVDCT_G5	39.78	0.01	0.07	13.91	0.19	44.95	0.23	0.31	99.45	85.03
FMVDCT_H5a	39.75	0.01	0.06	14.37	0.23	44.73	0.24	0.29	99.69	84.52
FMVDCT 15	38.84	0.03	0.04	18.18	0.26	41.51	0.25	0.16	99.28	80.05
FMVDCT .I5	38.36	0.03	0.02	21.74	0.38	38.50	0.28	0.09	99.40	75.63
FMVDCT A6	38.65	0.02	0.05	19.76	0.30	40.07	0.20	0.09	99.13	78.08
EMVDCT B6	39.09	0.01	0.05	17.09	0.27	42.36	0.20	0.19	99.30	81-31
	00 00	001	5.05	17 03	021	72 30	024	0.13	55 50	01:01

Table A1: Olivine core compositions measured by EPMA

Table A1: Continued

Crystal	Oxide (w	t %)								Mg#
	SiO ₂	TiO ₂	Al ₂ O ₃	FeO	MnO	MgO	CaO	NiO	Total	
FMVDCT_C6	39·16	0.02	0.04	16.16	0.25	42·87	0.25	0.21	98·95	82·32
FMVDCTD6	38·50	0.03	0.02	19.63	0.31	40.14	0.27	0.13	99.02	78·21
FMVDCTE6	38·23	0.02	0.00	21.59	0.34	38.44	0.26	0.08	98.97	75·76
FMVDCTF6	39.78	0.03	0.04	14·52	0.20	44·67	0.21	0.31	99.75	84·40
FMVDCTG6	38·38	0.04	0.02	20.53	0.32	39.54	0.26	0.12	99·21	77·17
FMVDCTH6	38·96	0.02	0.04	18·10	0.28	41·58	0.27	0.13	99.38	80.13
FMVDCT16	39.02	0.02	0.04	17·55	0.26	42·31	0.27	0.22	99.68	80.89
FMVDCTJ6	39·37	0.01	0.04	18·50	0.30	41.38	0.26	0.11	99.97	79·69
FMVDCTA7	38·32	0.02	0.02	22·11	0.34	38.88	0.26	0.02	100.01	75·53
FMVDCT B7a	38·85	0.04	0.03	19.36	0.34	40.55	0.25	0.17	99.59	78·58
FMVDCT C7	38.63	0.02	0.02	20.95	0.30	39.54	0.28	0.13	99.86	76·84
FMVDCT D7	38·81	0.03	0.03	18.99	0.28	41.16	0.26	0.14	99.69	79·20
FMVDCT E7	40.05	0.02	0.06	13.45	0.19	45.78	0.25	0.37	100.16	85.68
FMVDCT F7	38·81	0.02	0.02	20.15	0.32	40.10	0.27	0.17	99.85	77.74
FMVDCT G7	40.26	0.02	0.05	11.86	0.18	46.72	0.23	0.34	99.64	87.37
EMVDCT H7	38.86	0.03	0.04	19.70	0.30	40.30	0.28	0.11	99.61	78.22
FMVDCT I7	39.10	0.00	0.05	18-18	0.26	40.00	0.25	0.15	99.78	80.16
FMVDCT I7	39.37	0.01	0.06	16.06	0.20	43.77	0.23	0.23	99.94	82.74
EMVDCT A8	38.68	0.03	0.04	20.25	0.28	40.19	0.28	0.15	99.88	77.73
EMVDCT B8	30.40	0.02	0.04	17.44	0.26	42.50	0.25	0.26	100.26	91.06
	20.12	0.02	0.03	19.24	0.20	42.50	0.29	0.10	99.95	80.00
EMVDCT D8	30.28	0.02	0.05	17.12	0.27	42.63	0.25	0.10	99.80	91.29
EMVDCT E8	38.93	0.04	0.04	10.37	0.27	40.70	0.28	0.11	99.77	78.69
	20.21	0.02	0.02	17.00	0.20	40.70	0.25	0.19	100.06	91.60
EMVDCT_08	30.08	0.03	0.03	19.22	0.28	42.55	0.25	0.17	99.79	79.95
	39.75	0.02	0.02	10.52	0.20	40.67	0.27	0.12	99.86	79.39
	20.44	0.02	0.04	15.07	0.22	40.07	0.26	0.22	99.00	20.20
	20.60	0.02	0.04	14.47	0.10	43.37	0.20	0.22	99.54	02.03
	39.00	0.00	0.05	14.47	0.19	44.09	0.24	0.33	99.55	04.40
FINIVDCT_A9	39.52	0.01	0.02	14.01	0.20	44.74	0.25	0.30	99.67	04.34
FMIVDCTB9	39.27	0.01	0.03	10.07	0.25	43.20	0.25	0.23	99.81	82.07
FMVDC1_C9	39.05	0.01	0.04	17.72	0.27	42.16	0.26	0.20	99.73	80.69
FMVDCT_D9	39.20	0.02	0.05	17-19	0.25	42.57	0.21	0.19	99.69	81.31
FMIVDCT_E9	39.50	0.02	0.04	16.90	0.23	42.84	0.22	0.23	99.99	81.68
FMIVDC1_F9	38.30	0.02	0.03	21.07	0.31	39.00	0.28	0.09	99-10	76.48
FMVDC1G9	38.78	0.03	0.03	20.92	0.28	39.50	0.30	0.03	99.86	/6.85
FMVDC1_H9	39.21	0.02	0.02	16.99	0.58	42.51	0.26	0.21	99.50	81.44
FMVDCT19	39·18	0.05	0.04	16.94	0.24	42.62	0.53	0.53	99.49	81·56
FMVDCT_J9	39.78	0.01	0.02	13.64	0.19	45·22	0.24	0.29	99.42	85.35
FMVDCT_A10	39.70	0.05	0.02	14.19	0.17	44.85	0.53	0.32	99.52	84·77
FMVDCT_B10	38.94	0.02	0.02	18·76	0.27	41.56	0.21	0.19	100.01	79·56
FMVDCT_C10	39.36	0.01	0.02	18.65	0.26	41.76	0.53	0.12	100.43	79·75
FMVDCT_D10	39.20	0.03	0.04	18·01	0.27	41.82	0.56	0.19	99.82	80·31
FMVDCT_E10	39.02	0.03	0.02	18·13	0.24	41.67	0.24	0.22	99.60	80.17
FMVDCT_F10	39.28	0.03	0.02	16.13	0.23	43.45	0.26	0.58	99.69	82·56
FMVDCT_G10	38·55	0.01	0.02	20.82	0.32	39.59	0.59	0.10	99·71	76·95
FMVDCT_H10	39.07	0.01	0.06	18·12	0.26	41·75	0.53	0.15	99.64	80·20
FMVDCT_I10	39.42	0.01	0.02	16.06	0.25	43.38	0.27	0.30	99.74	82·58
FMVDCT_J10	38.41	0.03	0.03	22.20	0.36	38.42	0.24	0.04	99.74	75·22

Electron probe microanalysis data for 100 olivine crystal cores from a sample of Fimmvörðuháls tephra, erupted in 2010. Crystals were mounted in resin on a 10×10 grid, labelled A-J, 1-10. A 15 kV beam at 30 nA was used. Elemental abundances were measured against known standards using peak (P) and background (B) counting times as follows: Si P=20s, B=10s; Ni P=50s, B=25s; Fe P=30s, B=15s; Mn P=10s, B=5s; Al P=60s, B=30s; Ca P=50s, B=25s; Ti P=20s, B=10s; Mg P=20s, B=10s. Mg#=100 Mg/[Mg+Fe(t)+Mn]. Internal standards used were olivine and hematite.

Table	<i>e A2</i> :	XMT	averaged	grey-
scale	values	of cryst	al core sp	ots

Crystal	Average greyscale value
1	22036·62
2	22240.40
3	22304·52
4	22318·63
5	22332·52
6	22386.92
7	22386.94
8	22411.09
9	22428.06
10	22497.01
11	22500·18
12	22532·20
13	22563·31
14	22599.51
15	22644.19
16	22680.95
17	22693.69
18	22701-87
19	22717:55
20	22755-27
20	22755-69
21	22753 03
22	22702 32
23	22770.07
24	22770195
20	22791.00
20	22003-03
27	22020-04
20	22007.00
29	22872.01
30	22879-71
31	22886.65
32	22889-91
33	22897.38
34	22898-13
35	22902.99
36	22919.73
37	22920.37
38	22927.86
39	22943·27
40	22956.69
41	22964.50
42	22970.34
43	22971.84
44	22977-42
45	22982.00
46	22985.00

Crystal	Average grey	vscale valu
47	23000.41	
48	23002.69	
49	23002.72	
50	23007.99	
51	23012·22	
52	23013.74	
53	23017·99	
54	23038·58	
55	23047.08	
56	23052·95	
57	23063.83	
58	23069.55	
59	23069.61	
60	23069.73	
61	23089.71	
62	23089.94	
63	23107.06	
64	23127·45	
65	23132·32	
66	23133·42	
67	23136·40	
68	23149.36	
69	23152·86	
70	23173.85	
71	23183·52	
72	23210.50	
73	23211·85	
74	23212·57	
75	23217.59	
76	23230.23	
77	23254.65	
78	23260.12	
79	23310.84	
80	23425.39	
81	23444·72	
82	23456.44	
83	23467.67	
84	23475.56	
85	23476.79	
86	23494.71	
87	23518·67	

X-ray micro-computed tomography data from the spots placed on 87 olivine crystal cores from a sample of Fimmvörðuháls tephra, erupted in 2010, calculated using the workflow detailed in the text and summarized in Fig. 2. Each spot is approximately 25 μ m in diameter and 1 voxel deep.

EPMA data			XMT data					
Mg#	Rank	Linear range	Differential*	Average greyscale value	Rank	Linear range	Differential*	
75·02	100	75·02	0.00	22036·62	87	22037	0	
75·22	99	75·15	0.07	22240.40	86	22054	-187	
75·53	98	75·27	0.26	22304·52	85	22071	-233	
75·63	97	75·40	0.23	22318.63	84	22088	-230	
75·76	96	75·52	0.23	22332·52	83	22106	-227	
75·96	95	75·65	0.31	22386.92	82	22123	-264	
76·00	94	75·77	0.23	22386.94	81	22140	-247	
76·07	93	75·90	0·17	22411.09	80	22157	-254	
76·25	92	76·02	0.23	22428.06	79	22174	-254	
76·48	91	76·15	0.33	22497.01	78	22192	-305	
76·79	90	76·27	0.51	22500.18	77	22209	-291	
76.80	89	76·40	0.40	22532·20	76	22226	-306	
76.84	88	76·52	0.32	22563·31	75	22243	-320	
76.85	87	76·64	0.21	22599.51	74	22261	-339	
76.89	86	76·77	0.12	22644·19	73	22278	-366	
76·95	85	76·89	0.06	22680.95	72	22295	-386	
77·17	84	77·02	0.15	22693.69	71	22312	-381	
77·23	83	77·14	0.09	22701·87	70	22330	-372	
77·73	82	77·27	0.46	22717·55	69	22347	-371	
77.74	81	77.39	0.35	22755·27	68	22364	-391	
78·08	80	77·52	0.56	22755.69	67	22381	-374	
78·21	79	77.64	0.56	22762·92	66	22399	-364	
78·22	78	77.77	0.45	22770·07	65	22416	-354	
78·31	77	77.89	0.42	22770·12	64	22433	-337	
78·38	76	78·02	0.36	22791·85	63	22450	-342	
78·58	75	78·14	0.44	22803.65	62	22467	-336	
78·61	74	78·27	0.34	22826.04	61	22485	-341	
78·66	73	78.39	0.27	22867.06	60	22502	-365	
78·69	72	78·52	0·17	22872·01	59	22519	-353	
78·77	71	78.64	0.13	22879.71	58	22536	-343	
79·20	70	78·77	0.44	22886.65	57	22554	-333	
79·21	69	78·89	0.32	22889·91	56	22571	-319	
79·43	68	79·01	0.42	22897.38	55	22588	-309	
79·56	67	79·14	0.42	22898·13	54	22605	-293	
79·67	66	79·26	0.41	22902.99	53	22623	-280	
79·69	65	79·39	0.30	22919.73	52	22640	-280	
79·74	64	79·51	0.23	22920.37	51	22657	-263	
79·75	63	79·64	0·11	22927.86	50	22674	-254	
79·75	62	79·76	-0.02	22943.27	49	22691	-252	
79·92	61	79·89	0.04	22956.69	48	22709	-248	
79·95	60	80·01	-0.06	22964.50	47	22726	-239	
80.00	59	80·14	-0·14	22970·34	46	22743	-227	
80.05	58	80·26	-0·21	22971.84	45	22760	-211	
80.11	57	80.39	-0.28	22977-42	44	22778	-200	
80.13	56	80.51	-0.39	22982.00	43	22795	-187	
	50							

Table A3: Rank-order and linear-array order of EPMA and XMT data

Table A3: Continued

EPMA data				XMT data					
Mg#	Rank	Linear range	Differential*	Average greyscale value	Rank	Linear range	Differential*		
80·13	55	80.64	-0·51	22985.00	42	22812	-173		
80·16	54	80.76	-0.60	23000.41	41	22829	-171		
80·16	53	80.89	-0.73	23002·69	40	22847	-156		
80·17	52	81·01	-0·84	23002·72	39	22864	-139		
80·19	51	81·13	-0·95	23007.99	38	22881	-127		
80·20	50	81·26	-1.06	23012·22	37	22898	-114		
80·31	49	81·38	-1.08	23013.74	36	22916	-98		
80.40	48	81·51	-1·11	23017.99	35	22933	-85		
80.49	47	81.63	-1·15	23038·58	34	22950	-89		
80·55	46	81.76	-1·21	23047.08	33	22967	-80		
80.69	45	81.88	-1.20	23052.95	32	22984	-69		
80·89	44	82·01	-1·11	23063·83	31	23002	-62		
81.06	43	82·13	-1·07	23069.55	30	23019	-51		
81·14	42	82·26	-1·12	23069.61	29	23036	-33		
81·28	41	82·38	-1.10	23069.73	28	23053	-16		
81·29	40	82·51	-1·21	23089.71	27	23071	-19		
81·31	39	82·63	-1·32	23089.94	26	23088	-2		
81·31	38	82·76	-1·44	23107.06	25	23105	-2		
81·38	37	82·88	-1.50	23127.45	24	23122	-5		
81.38	36	83·01	-1.62	23132-32	23	23140	7		
81.43	35	83·13	-1.70	23133-42	22	23157	23		
81.44	34	83.25	-1.82	23136.40	21	23174	38		
81.50	33	83.38	-1.88	23149:36	20	23191	42		
81.56	32	83.50	-1.95	23152.86	19	23208	56		
81.60	31	83.63	-2.03	23173.85	18	23226	52		
81.68	30	83.75	-2.08	23183-52	17	23243	59		
81.78	29	83-88	-2.10	23210.50	16	23260	50		
81.90	28	84.00	-2.11	23211.85	15	23277	66		
82.07	20	84.13	-2:06	23217-05	14	23295	82		
82.32	26	84.25	_1.93	23212.57	13	23233	94		
82.56	20	84.38	-1.82	23230.23	12	23379	99		
82.58	20	84.50	_1.92	23254.65	11	23346	92		
02.50	24	94.62	2.04	22224 00	10	20040	102		
02.03	20	84.75	-2.04	23200.12	0	23304	70		
02.03	22	04.75	-2.07	23310:04	9	23301	70		
02.74	21	85 00	-2.13	23423-33	0	23330	-27		
02.00	20	00 ^{.00}	- 1.24	23444.72	6	23415	-29		
03.00	19	00.12	-1.24	23430'44	5	23433	-24		
04.31	10	05.25	-0.94	23407.07	5	23450	- 10		
84.34	17	85.38	-1.03	23475.50	4	23467	-9		
04.37	10	00.00	- 1.13	234/0./3	3	23484	/		
84·40	15	85.0∠ 05.35	- 1.22	23494.71	2	23501	/		
84.46	14	85.75	-1.29	23518.67	1	23519	0		
84.52	13	85.87	-1:35						
84·77	12	86.00	-1.23						
84·83	11	86·12	-1·29						

EPMA data				XMT data					
Mg#	Rank	Linear range	Differential*	Average greyscale value	Rank	Linear range	Differential*		
85·03	10	86·25	-1·22						
85.08	9	86·37	-1·30						
85·17	8	86.50	-1·33						
85·35	7	86.62	-1·27						
85·38	6	86·75	-1·37						
85.68	5	86.87	-1·19						
86·41	4	87.00	-0·58						
86·81	3	87·12	-0·31						
87·10	2	87·25	-0·14						
87·37	1	87.37	0.00						

Rank-order and linear-array-order data used to generate plots displayed in Fig. A1. It should be noted that higher greyscale values denote lighter voxels = lower Mg#. The XMT linear-array order is inverted to allow comparison against Mg#. Four significant figures are retained when describing EPMA data, to maintain unique numbers and illustrate the shape of the arrays.

*The differential captures the degree to which each data point (each crystal) departs from a straight line between the highest and lowest value on the corresponding rank-order plot, so that comparisons are easier to make visually (see Fig. A1).

Bin no.	Range	
	from	to
1	23550	+
2	23378	23550
3	23206	23378
4	23033	23206
5	22861	23033
6	22689	22861
7	22517	22689
8	22344	22517
9	22172	22344
10	22000	22172
11	-	22000

Table	<i>A4</i> :	Bins	used	to	generate	frequency	diagram	from
XMT	grey.	scale s _i	pots					